

double/triple integrals

2D: double integrals

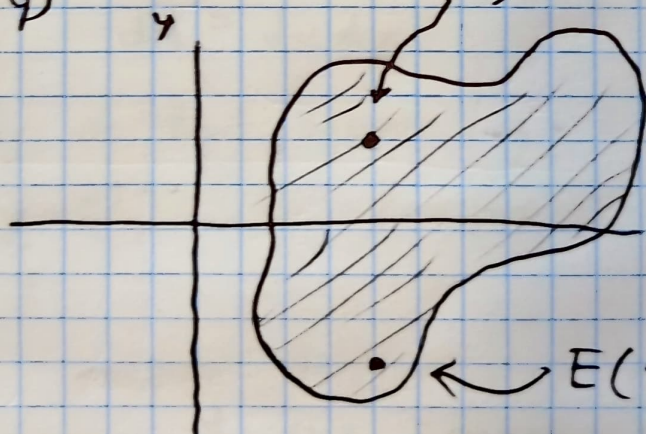
Suppose we have some region in x, y -land that we wish to integrate over

For example, we may wish to find the total energy of a region of space where the energy density is defined at every point (x, y) i.e.

Energy density is a function of (x, y)

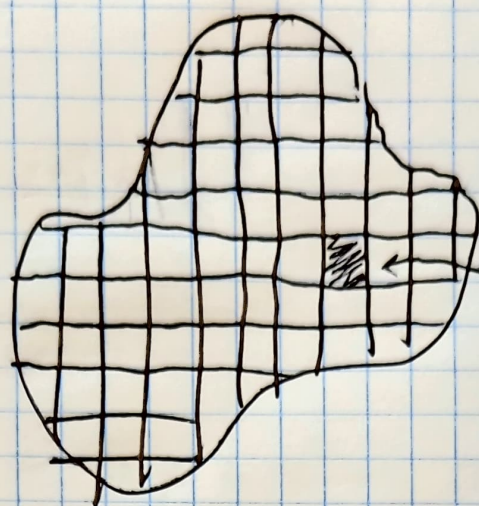
$$E = E(x, y)$$

$$E(4, 2) = 7\text{J}$$



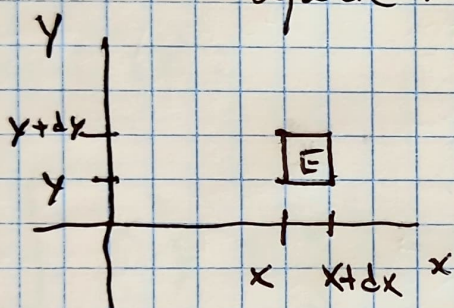
$$E(4, -3) = 3\text{J}$$

The best way to do this is to break the region up into little squares, find the energy for each little square & add them up.



one
arbitrary
square

we ask "what is the area of one small square?"



geometry tells us $dA = dx dy$

the energy density of that square is $E(x, y)$

or $E(x + \frac{dx}{2}, y + \frac{dy}{2})$ etc (turns out $E(x, y)$ is just dandy)

So the total energy of one square is

$$E(x, y) dA = E(x, y) dx dy$$

Now we add them up & take $dx, dy \rightarrow 0$

$$E_{Total} = \int_{y_i}^{y_f} \int_{x_i}^{x_f} E(x,y) dx dy \quad (\text{for a rectangular region})$$

E.G. $E(x,y) = xy + x^2 - y^3$

$$E_T = \int \int (xy + x^2 - y^3) dx dy$$

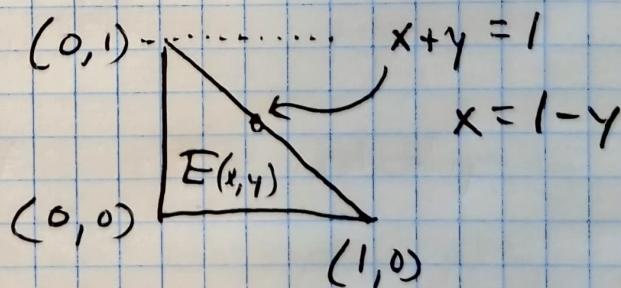
inside integral (st):

$$= \int \frac{x^2}{2} y + \frac{x^3}{3} - y^3 x dy$$

$$= \frac{x^2 y^2}{4} + \frac{x^3 y}{3} - \frac{y^4 x}{4} + \text{constant}$$

If our region is not rectangular then our integration limits may be functions

E.G. a triangle



$$\int_0^1 \int_0^{1-y} E(x,y) dx dy$$

in the x-integral $x=1-y$

we can also find average values over a region. Just integrate & divide by the total area.

$$\text{Area} \equiv \iint dA = \iint dx dy$$

e.g. area of a rectangle

$$\begin{aligned} A_R &= \int_0^b \int_0^a dx dy = \int_0^b x \Big|_0^a dy \\ &= \int_0^b a dy = \boxed{ab} \end{aligned}$$

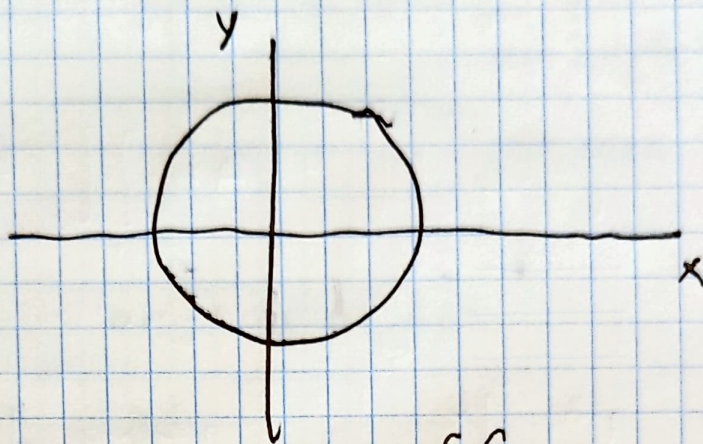
$$\langle E \rangle = \frac{\int E(x,y) dx dy}{\int dx dy}$$

$$\langle E^2 \rangle = \frac{\int E^2(x,y) dx dy}{\int dx dy}$$

etc

Often physical problems are rotationally symmetric & cartesian coordinates are cumbersome compared to polar

Suppose we have a circular region



we could again do this via $\iint dx dy$
but the integration limits are weird thanks

to

$$x^2 + y^2 = R^2$$

or

$$x^2 = R^2 - y^2 \rightarrow \pm \sqrt{R^2 - y^2}$$

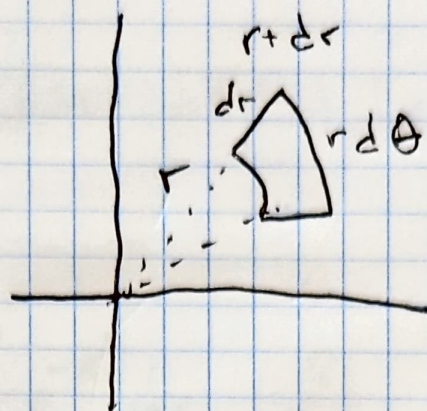
is it's a mess

but in ~~for~~ polar coordinates describing a
circle is easy!

$$r = R$$

$$0 \leq \theta \leq 2\pi$$

so let's chop our region up using polar
coordinates.



$$\text{arc length} = r d\theta$$

$$\Rightarrow \text{area} = r dr d\theta$$

$$dA = r d\theta dr \quad \text{or} \quad r dr d\theta$$

the

$$A_{\text{circle}} = \int_0^{2\pi} \int_0^R r dr d\theta = 2\pi \frac{R^2}{2} = \pi R^2$$

you could also go a different route

$$x = r \cos \theta$$

$$y = r \sin \theta$$

$$dx = d(r \cos \theta) = dr \cos \theta - r \sin \theta d\theta$$

$$dy = d(r \sin \theta) = dr \sin \theta + r \cos \theta d\theta$$

then multiply $dx dy = r dr d\theta$

★ (check this) ★

* A MAJORLY IMPORTANT EXAMPLE *

How do we compute

$$I = \int e^{-x^2} dx \quad ?$$

we can't use a substitution or integrate by parts but we can change coordinates to polar

Trick consider I^2

$$I^2 = \int e^{-x^2} dx \int e^{-y^2} dy$$

$$= \iint e^{-(x^2+y^2)} dx dy$$

A-HA! $x^2 + y^2$ looks like a circle!

let's change to polar coords!

$$r^2 = x^2 + y^2$$

$$dx dy = r dr d\theta$$

initially we seem stuck.

were taking e^{-x^2} for e^{-r^2} which looks the same, but the r in $r dr d\theta = dA$ saves the day!

$$I^2 = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{-(x^2+y^2)} dx dy = \int_0^{2\pi} \int_0^{\infty} e^{-r^2} r dr d\theta$$

notice the limits, when we switch to r we go from $0 \rightarrow \infty$; not $-\infty \rightarrow \infty$

the θ integral is easy $\int_0^{2\pi} d\theta = 2\pi$

$$\text{so } I^2 = 2\pi \int_0^{\infty} e^{-r^2} r dr$$

$$\text{let } u = -r^2 \Rightarrow du = -2r dr$$

$$\Rightarrow 2r dr = -\frac{du}{2}$$

$$\Rightarrow I^2 = \frac{2\pi}{2} \int_{-\infty}^0 e^u du = \pi ! \text{ WOW!}$$

$$\Rightarrow \boxed{I = \sqrt{\pi} = \int_{-\infty}^{\infty} e^{-x^2} dx}$$

* HOME WORK *

solve
$$I = \int e^{-ax^2} dx$$

you'll find
$$I = \sqrt{\frac{\pi}{a}}$$

once we have
$$\int e^{-ax^2} dx$$

we can have BIG Fun

* for example *

$$\int x^2 e^{-ax^2} dx$$

doesn't $x^2 e^{-ax^2}$ look a lot like

$$-\frac{\partial}{\partial a} e^{-ax^2} ?$$

then
$$\int x^2 e^{-ax^2} dx = - \int \frac{\partial}{\partial a} (e^{-ax^2}) dx$$

since we're not integrating over a we can pull $\frac{\partial}{\partial a}$ out of integral!

$$\begin{aligned}
 \Rightarrow \int x^2 e^{-ax^2} dx &= -\frac{\partial}{\partial a} \int e^{-ax^2} dx \\
 &= -\frac{\partial}{\partial a} \left(\sqrt{\frac{\pi}{a}} \right) = -\sqrt{\pi} \left(\frac{\partial}{\partial a} a^{-\frac{1}{2}} \right) \\
 &= \frac{\sqrt{\pi}}{2} a^{-3/2} \\
 &= \boxed{\sqrt{\frac{\pi}{a}} \frac{1}{2a}}
 \end{aligned}$$

recall that

$\langle x \rangle = \int x f(x) dx$ is the average of x over some space.

$$\frac{1}{2a} \sqrt{\frac{\pi}{a}} = \langle x^2 \rangle = \int x^2 f(x) dx$$

where $f(x)$ is e^{-ax^2}

this special $f(x) = e^{-ax^2}$ is what's known as the "Gaussian distribution"

or "Bell Curve" or "Normal Distribution"

In its full glory

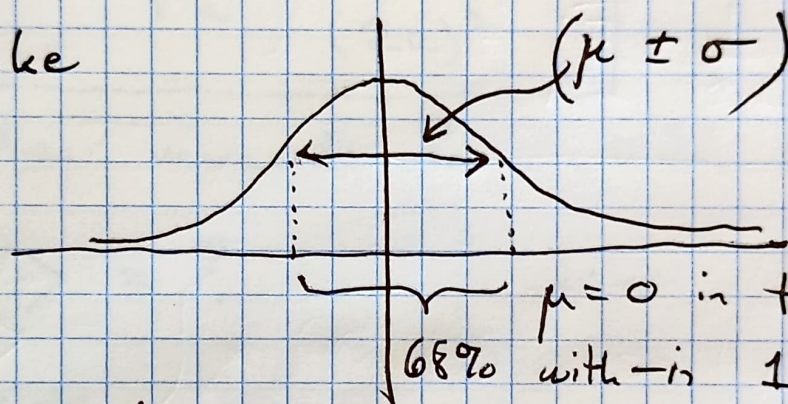
$$p(x) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{1}{2} \left(\frac{x-\mu}{\sigma} \right)^2}$$

where

$\mu \equiv$ mean

$\sigma \equiv$ standard deviation

it looks like



$$\int_{-\infty}^{\infty} p(x) dx = 1$$

HOMEWORK

check the following

$$\mu = \langle x \rangle = \int_{-\infty}^{\infty} x p(x) dx$$

$$1 = \int_{-\infty}^{\infty} p(x) dx$$

σ^2

we get a nice class of integrals with this trick called the "moments" of $p(x)$

$$\boxed{\begin{aligned}\langle x^{2n} \rangle &= \int_{-\infty}^{\infty} x^{2n} e^{-ax^2} dx \\ &= \sqrt{\frac{\pi}{a}} \frac{(2n-1)!!}{(2a)^n}\end{aligned}}$$

what are the moments for odd powers?

$$\langle x^{(2n+1)} \rangle = ?$$

★ HOMEWORK ★

Show $\langle x^{2n+1} \rangle = 0$ for
Gaussian (take $\mu=0$ & $\sigma=1$)

★ Hint ★ use a trick you recently learned!

★ Think, then calculate ★

★ PHYSICS ★

Recall that for a box of gas the temperature is related to the average velocity of the molecules via

$$kT = \frac{1}{2} m \langle v_x^2 \rangle$$

in each dimension

In a 3D Box we have

$$kT = \frac{3}{2} m \langle v^2 \rangle$$

where $\langle v^2 \rangle = \langle v_x^2 + v_y^2 + v_z^2 \rangle = v_{rms}^2$
 $v_{rms} = \sqrt{\langle v^2 \rangle}$

You might ask: "what probability function or distribution gives rise to this?"

Maxwell figured out this distribution of speeds is given by

$$f(v_x) = \sqrt{\frac{m}{2\pi kT}} e^{-\frac{mv_x^2}{2kT}}$$

a Gaussian in v_x !

$f(v_x)$ tells us what fraction of molecules in our box have velocity v_x at some temperature T .

e.g. for small T we would expect the fraction of molecules moving crazy fast to be roughly zero

★ Check this ★ Taylor expand $f(v_x)$ for $T \ll 1$ as a function of T

Now, because $f(v_x)$ represents a fraction of a total we demand

$$\int_{-\infty}^{\infty} f(v_x) dv_x = 1$$

This demand determines the $\sqrt{\frac{m}{2\pi kT}}$ factor in front.

Well it's Gaussian so we can find the average SPEED. Emphasis on "Speed"
the average "velocity" will be ZERO

due to symmetry [odd · even $\rightarrow 0$]

$$\bar{s} = \langle |v_x| \rangle = \int_0^\infty 2 v_x f(v_x) dv_x$$

Notice the odd/even trick doesn't work because $|v_x|$ is even & $f(v_x)$ is even

$$\langle v_x \rangle = 0$$

$$\text{but } \langle |v_x| \rangle \neq 0$$

you can do this integral (TRY)
≡ find that

$$\boxed{\bar{s} = \langle |v_x| \rangle = \sqrt{\frac{2kT}{\pi m}}}$$

This has all been in 1D
for 3D we have

$$v^2 = v_x^2 + v_y^2 + v_z^2$$

$$\therefore \boxed{f(v) = 4\pi^2 v^2 \left[\frac{m}{2\pi kT} \right]^{3/2} e^{-\frac{mv^2}{2kT}}}$$

Notice the v^2 term in front now.

this term comes from "spherical coordinates"

$$\int dx dy dz \rightarrow \int v^2 \sin \theta dv d\theta d\phi$$

similar to changing from cartesian to polar

Polar

$$x = r \cos \theta$$

$$y = r \sin \theta$$

$$dA = dx dy = r dr d\theta$$

SPHERICAL

$$x = r \cos \theta \sin \phi$$

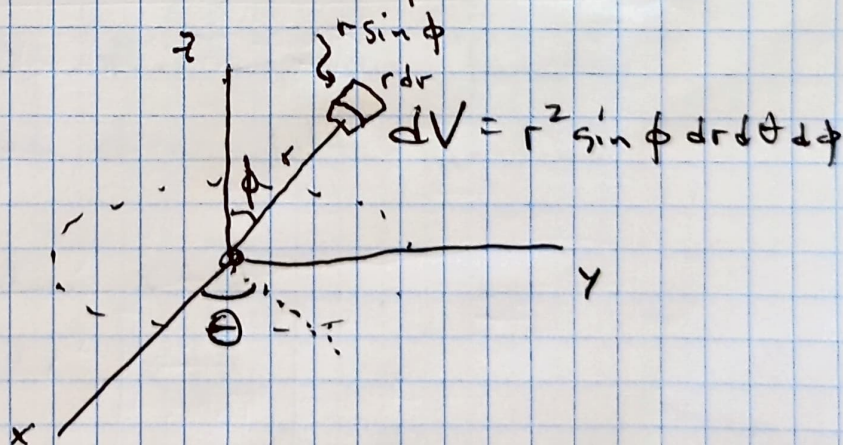
$$y = r \sin \theta \sin \phi$$

$$z = r \cos \phi$$

$$dV = dx dy dz = r^2 \sin \theta dr d\theta d\phi$$

ϕ : azimuthal angle

θ : polar angle

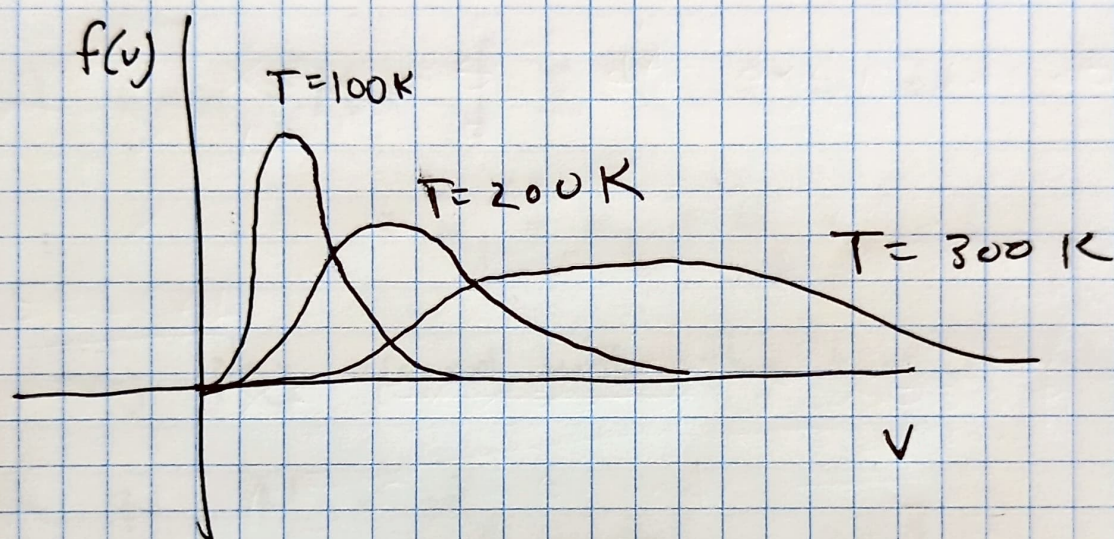


Someday you'll memorize this; especially $r^2 \sin \theta dr d\theta d\phi$

$$V_{sp} = \frac{4}{3} \pi R^3 = \int_0^{2\pi} \int_0^{\pi} \int_0^R r^2 \sin \theta dr d\theta d\phi$$

you won't need to use spherical coordinates yet but that is why $f(v_k) \neq \tilde{f}(v)$ have different forms

$f(v)$ looks like



as T grows the standard deviation increases
as $T \rightarrow 0$ we expect $f(v) = 1$ for $v=0$
& $f(v)=0$ for $v \neq 0$ this specific "function"
has a name "the dirac-delta-function"

$$\delta(x) = 1 \quad \text{when } x=0$$
$$= 0 \quad \text{when } x \neq 0$$

$\delta(x)$ allows us to do useful things like
if you have a bunch of point masses
we can write their density as $\rho(x) = \sum m_i \delta(x-x_i)$

usually we integrate $\delta(x)$ ~~age~~ against
a function

$$\int \delta(x-a) f(x) dx = f(a)$$

which ends up being SUPER handy.

$$\begin{aligned} \text{total mass} = M &= \int \sum \delta(x-x_i) m_i dx \\ &= m_1 + m_2 + m_3 + \dots \end{aligned}$$

we may play around with this further later.

Back to Maxwell,

$$f(v) = 4\pi \left(\frac{m}{2\pi kT} \right)^{3/2} v^2 e^{-\frac{mv^2}{2kT}}$$

find $\langle |v| \rangle$ = average speed

$$\langle |v| \rangle = 2 \int_0^{\infty} v f(v) dv$$

using our
Gaussian integral
from earlier.